- Notice that the plastic disk with the attached diffraction grating can be rotated. Looking in the spectrometer, you should notice that the colors move as you rotate the disk. Botate the disk until you see colors in a horizontal line to your left. The colors should appear between the two lines of numbers on the scales.
- 3. Look at an illuminated incandescent bulb through the spectremeter. Be careful to aim the slit (on the right side of the spectrometer) at the light bulb and look straight ahead at the spectrum on the scale. You should see a continuous spectrum of colors from red through violet. On the scale below (Figure 2), draw the colors you see. Use colored pencils to shade in the observed colors. The various colors are described by wavelength in nanometers (nm) or by the energy of the particle of light, a photon, expressed in electron-Volts (eV). (1 eV = 1.6022 X 10-19 joule)

ſ	Company of the contract of the			
l	12 12 14 14 24 1 127	26 22 20 37 34 3 1001 3 1 10000	eV	
ļ	1 1 1 1 1 1 1 1 1 1	1 200 11 11 111		
l				
l				
١	1 2.3 000 04	[11 5] 12 (12 15 11)	am	į
i	. 500	. 550 400		
i				sLit

Figure 2

- Read the number on the scale corresponding to the light farmest to the right that you can see and the number corresponding to the light farthest to the left that you can see.
- a) The observed spectrum extends from _____nm to _____nm.
- b) The colors at mase places on the scale are:

5. Now look at a fluorescent light through the spectrometer
Describe the spectrum you see. Is it different from the
spectrum you observed in Step 1?

6.	Again record the ends	of the spectrum.	
	The colored spectrum	extends	
	from	nm to	nm

The spectrum from the fluorescent light should include several bright vertical "lines". These are images of the slit. Indicate the positions of these lines on the scale below.

17 14 17 26 17 24 17 24 17 25 27 47 47 47 47 47 47 47 47 47 47 47 47 47	
	thi

Figure 3

7. Read the positions of the bright lines on the scale and record them in Table 1.

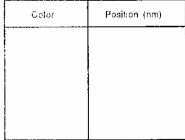


Table 1

8. The most common type of fluorescent light will have the mercury emission lines superimposed on a continuous spectrum. The scale should be adjusted so that the bright green line in the spectrum of the fluorescent light is at 546 nm (notice ine small mark on the scale). Adjust the scale by insering the tip of a pen or pencil into one of the sprocket holes on the time, (see Figure 1).

NOTE: If the green line is to the LEFT of the 546 nm mark, push to film to the right (AS SEEN THROUGH THE SPECTROMETER). If the green line is to the right of the 546 nm mark, move the film to the LEFT.

 Point the slit of your spectrometer at a white surface that has fluorescent light shining on it, such as a wall or a movie screen, and measure the ends of the spectrum and the positions of any bright lines that you see. Record your data in Table 2.

1, 1, 1, 2, 2, 2, 1, 1, 2, 2, 1, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2, 2,
Position (nm)

Table 2

- a) Compare the results of Steps 5 and 6. Was the spectrum that you saw from the fluorescent light similar to or different from the spectrum you saw when you looked at the white surface?
- b) Why do you think the spectra were similar or different?

10. Use your spectrometer to observe as many other light sources as you can find. These could include the red or green LEDs (Light Emitting Diodes) on a VCR (Video Cassette Recorder) or stereo system; chemical light sticks; and "neon" signs in the windows of stores and restaurants. List the object and describe the spectrum you observed. (Are there any bright or dark lines in the spectrum? If there are any bright or dark lines, give the positions and the colors of the lines.)

Object 1:

Color as seen without spectrometer:

Description of spectrum:

Object 2:

Color as seen without spectrometer:

Description of spectrum:

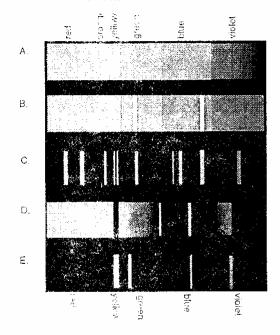
Object 3:

Color as seen without spectrometer:

Description of spectrum:

12. Observe the spectra of some indoor lights and a few street lights or security lights. White the letter of the observed spectrum (on the left) on the blank that corresponds with the type of light (and its apparent color) below.

AL FOTEIA OF SELECTED LIGHTS



incandescent light bolb	Liga pressur e s odium
fluorest ent light	low pressure sodium
m	PECHEZ VADOR

page 6 w1993 Learning Technologies, Inc.

13.	Take your spectrometer outside (better yet, look through a
	window from a darkened room) and point the slit toward the
	bright sky near the Sun.

oright sky hear the Sun.

DO NOT LOOK GIRECTLY AT SUM!!

IT GAN DAMAGE YOUR EYES!!

You should see a spectrum of all the colors with narrow, dark lines superimposed.

a)	The spectrum extends		i		
	from		nm to		٦m

Now measure the positions of the dark lines that you see. Record the results in Table 3.

Missing Color	Position (nm)

Table 3

- Compare these absorption lines in the Sun's spectrum with those listed in Table 4.
- a) What elements do you conclude are present in the Sun?
- b) Do you think that you have found all the elements that are in the Sun? Why or why not?
- c) Where do you expect that elements would have to be located in order to cause dark absorption lines in the spectrum of the Sun? (Would they have to be located inside the Sun, or: the Sun's surface, above the Sun's surface, in space between the Sun and the Earth, or in the Earth's atmosphere?)

- 15. Point the spectrometer silt at a bright, white cloud.
- a) Describe the spectrum that you see. How does the "cloud" spectrum compare to the spectrum of the Sun? Does the cloud spectrum have dark lines as the solar spectrum does?
- b) Why do you think the cloud spectrum appears the way it does?
- 16. Look at the Moon through the spectroscope. (This activity is best done at night when the Moon is bright compared to the background sky, such as when there is a full Moon visible two or three hours after sunset. Also, the small apparent size of the Moon and the unsteadiness of holding the spectrometer will present a problem.)

Describe the spectrum. (How does the Moon's spectrum compare to the spectrum of the Sun? Does the lunar spectrum have dark lines as the solar spectrum does? Are they the same lines? Why or why not?)

ABSORPTION LINES IN THE SUN

(from the CRC Handbook of Chemistry and Physics)

Lines	Wavelength	Lines	Wavelength
due to	(nm)	due to	(nm)
Iron	372.8	Iron	516.8
Iron	382,0	Magnesium	516.7
Calcium	393.4	Magnesium	517.3
Calcium	396.8	Magnesium	518.4
Hydrogen	410.2	!ron	527.0
Calcium	422.7	Sodium	589.0
lron	430.8	Sodium	590.0
Hydrogen	434.0	Hydrogen	656.3
Hydrogen	486.1	Oxygen	759.4
		Oxygen	762.1

Table 4

OPTIONAL ACTIVITY

The purpose of this activity is to study how certain transparent materials will allow some colors of light through and absorb the others. Locate various pieces of colored transparent glass or cellophane or make a colored liquid. The liquid can be made by mixing food dyes and water in a clear glass or plastic container. (Don't forget good experimental practice: Check to see if the "clear" glass absorbs any colors before observing light through a colored liquid in the bottle AND observe the light source to make sure that it is producing all colors)

Place a transparent colored object (glass, cellophane, liquid) between a bright white light source and the spectrometer. For each object, record in Table 5 (on page 10) the following data: the type of object and its color, the missing color(s), and the position(s) of the dark bands on the spectrometer scale. (The dark bands, called absorption bands, are due to photons of certain wavelengths being absorbed by the object. When the photons are absorbed, the colors corresponding to the photon energies are removed from the spectrum and gaps, or bands, show up in the spectrum where the missing colors would have appeared if there was no absorption.)

NOTES

Object and Color	Missing Color	Position (nm)
.		

Table 5